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Complex environmental beta-plane turbulence: laboratory experiments with altimetric imaging velocimetry

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absence of β -effect) by Afanasyev and Craig (2013) gave the experimental evidence of dual cascade with the spectral slopes of $-5/3$ and -3 . Further experiments with barotropic turbulence on the β -plane by Zhang and Afanasyev (2014) demonstrated the dual cascade in the presence of β -effect as well as a “figure eight” energy spectrum in the wavenumber space.

Zonal jets have a long history of investigation starting from the pioneering experiments by Whitehead (1975) and Colin de Verdiere (1979). The jets readily form when a spatially localized forcing is applied in the β -plane fluid (e.g. Sommeria et al., 1988, 1989; Marcus and Lee, 1998; Afanasyev et al., 2011, 2012; Slavin and Afanasyev, 2012). However, a distributed forcing such as that provided by baroclinic instability, also creates jets. In the ocean, the regions where baroclinic instability is dynamically important include the Antarctic Circumpolar Current (ACC) as well as western boundary currents and their extensions. Multiple jets as well as mesoscale eddies are created there by the baroclinic instability thus forming a dynamically complex turbulent flow. A classic model for a baroclinically unstable system is a rotating annulus where heating/cooling is provided at the outer wall/center of the tank respectively. Different aspects of the dynamics of this system were studied in a number of experiments (Hide and Mason, 1975; Mason, 1975; Bastin and Read, 1997, 1998; Wordsworth et al., 2008; Smith et al., 2014). A somewhat different experimental approach to modelling a baroclinically unstable two-layer flow with vertical shear was used in a recent study by Matulka and Afanasyev (2015). It was shown that the meridional scale of the jets is determined to large extent by the radius of deformation and, at the same time, is in good agreement with the Rhines scale (Rhines, 1975). The jets are driven by (nonlinear) Reynolds stresses due to baroclinic meanders. This is in agreement with a scenario described by Berloff et al. (2009a, b). The authors describe the formation of jets as a secondary instability of the primary instability of the baroclinic flow in the form of mainly meridional motions (so-called “noodles”).

In this study we perform spectral analyses of the flows described in Matulka and Afanasyev (2015) and compare them with the results obtained for a somewhat different

where $f_0 = 2\Omega$ is the Coriolis parameter. The values of the β -parameter in our experiments were between 0.07 and $0.1 \text{ cm}^{-1} \text{ s}^{-1}$.

We use the Altimetric Imaging Velocimetry (AIV) system to measure the gradient of the surface elevation η (for more details see Afanasyev et al., 2009). AIV is based on optical altimetry first described in Rhines et al. (2006). Laboratory altimetry is not unlike the satellite altimetry which became an irreplaceable tool in oceanography. Apart from measuring $\nabla\eta$, AIV can be used as a tool to visualize the entire surface of rotating fluid just like satellite altimetry provides a global coverage of the Earth's oceans. The AIV also provides an alternative to a well-known Particle Imaging Velocimetry (PIV) technique. To obtain velocity field using PIV one has to find correlations between small areas (typically 12–48 pixels in both dimensions) of two successive images of the flow. If a camera has an imaging array of size say 1000×1000 pixels, the resulting array of velocity vectors has dimensions less than 100×100 . Thus, the PIV technique effectively reduces the spatial resolution by a factor of 100 or more. AIV, on the other hand, allows one to obtain the velocity vector in every pixel of the image, which makes its spatial resolution practically unlimited given that the cameras with large imaging arrays are now easily available. The limitation of the AIV is that it can only be used in a relatively fast rotating fluid with free surface. It, however, is perfectly suited for oceanographic fluid dynamics experiments on the β -plane as those described in this paper. Paraboloidal surface of the rotating fluid is used as a mirror of the Newtonian telescope. If the surface is disturbed by the pressure perturbations due to the flow, the slope of the surface changes slightly. These perturbations of the slope are detected by the AIV and measured using simple geometry and a color coding.

AIV measures an “exact” (within experimental accuracy) surface elevation gradient, $\nabla\eta$, which translates into the pressure gradient, $\nabla p = \rho g \nabla\eta$, at the surface. Note that the main uncertainty is due to the color noise of the imager. To reduce the noise we pass the data through a median filter with a window size of 5×5 pixels (physical size of approximately $0.2 \text{ cm} \times 0.2 \text{ cm}$). The overall uncertainty of $\nabla\eta$ can be estimated to

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be approximately 5%. The velocity field is not measured directly by the AIV but rather obtained from the measured $\nabla\eta$ using a (quasi-) geostrophic approximation.

Barotropic component of velocity can be calculated in geostrophic approximation as follows

$$5 \quad \mathbf{V} = \frac{g}{f_0} (\mathbf{n} \times \nabla\eta), \quad (3)$$

where \mathbf{n} is the vertical unit vector. A next order of approximation is provided by the quasi-geostrophy which gives

$$\mathbf{V} = \frac{g}{f_0} (\mathbf{n} \times \nabla\eta) - \frac{g}{f_0^2} \frac{\partial}{\partial t} \nabla\eta - \frac{g^2}{f_0^3} J(\eta, \nabla\eta), \quad (4)$$

10 where J is the Jacobian operator. The second and the third terms in the RHS of Eq. (4) are corrections to the geostrophic velocity which take into account transient and non-linear effects. Their relative importance is determined by the temporal Rossby number $Ro_T = 1/(f_0 T)$ and the Rossby number $Ro = U/(f_0 L)$ respectively. Here T is the time scale of the unsteady processes in the flow, while U and L are velocity and length scales of the flow. Thus, the velocity field is determined more accurately when the
 15 flow is closer to being quasigeostrophic. “Textbook” theory on the validity of the quasi-geostrophic approximation applies here. Since the Rossby number did not exceed unity even in the core of the eddies in our experiments and the mean values of the Rossby number were of the order of 10^{-1} , the velocity was, on average, within 10% of the “exact” velocity.

20 Relative vorticity, $\zeta = \mathbf{n} \cdot \text{curl}\mathbf{V}$, was calculated by differentiating the velocity field. Since numerical differentiation amplifies noise in the original data, we used the Sobel gradient operators with 5×5 kernels. The kernels are convolved with the velocity data to calculate the derivatives in x and y directions.

25 According to the Taylor–Proudman theorem, the surface velocity given by Eq. (4) is a good approximation for the velocity in the entire column of water except the Ekman

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experiment. Flows created by buoyancy sources in the rotating fluid (in the absence of the β -effect) were studied in laboratory experiments by Fernando et al. (1991) and Maxworthy and Narimusa (1994). The energy flux per unit area of the bottom of the tank is $Q = 2.5 \times 10^3 \text{ W m}^{-2}$ in our experiments. This translates into the buoyancy flux $B = \alpha g Q / c_p = 1.2 \times 10^{-6} \text{ m}^2 \text{ s}^{-3}$, where α is the thermal expansion coefficient and c_p is the volumetric heat capacity of water at constant pressure. A dimensionless parameter which controls the regime of the thermal convection is the Rayleigh number. The flux Rayleigh number is quite high in our experiments, $Ra_{\text{flux}} = BH_0^4 / \nu^2 \kappa = 1.8 \times 10^9$ (where κ is the thermal diffusivity of water) that indicates a regime of turbulent convection. Comparing to the recent experiments by Read et al. (2015) where similar heating was used but in an experiment of a larger scale, we note that our buoyancy flux was almost two orders of magnitude higher; the Rayleigh number was, however, an order of magnitude lower because of the smaller depth of water in our experiments.

Figure 3a and b shows the fields obtained as a result of the velocity calculation from measured $\nabla \eta$ in the thermal experiment. Azimuthal velocity, V_{az} in a polar coordinate system with the origin at the center of the tank is shown in panel a, while panel b shows the relative vorticity, $\zeta = \mathbf{n} \cdot \text{curl} \mathbf{V}$. The relative vorticity field shows a fine structure of the flow with multiple small scale cyclonic (red) and anticyclonic (blue) eddies. The relative vorticity is normalized by the Coriolis parameter f_0 , thus the image in panel b can be interpreted as a map of the Rossby number, $Ro = \zeta / f_0$. The values of Ro can reach unity in the strongest eddies while the rms value is approximately 0.2. An original alignment of filaments with the heating wires can still be seen in the central area of the tank while closer to the wall where the water is deeper and the β -effect is stronger, the bands of positive and negative vorticity are aligned in the zonal direction which indicates the presence of zonal jets. The jets can be seen more clearly in the azimuthal velocity image in panel a. The circulation is mainly in the counterclockwise (eastward) direction (V_{az} is positive, red color). Maximum values of V_{az} in the jets are approximately 1 cm s^{-1} .

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Figures 2b, 3c and d show the experiment with the saline forcing. The water in the tank is initially of salinity $S = 30$ ppt. When a source distributed along the wall of the tank delivers fresh water, it creates a current along the wall. This current is initially wedge-shaped in cross-section and is approximately in geostrophic balance such that it “leans” on the wall to its right. This coastal current can be seen clearly in Figs. 2b and 3c, the velocity of the current is initially in excess of 5 cm s^{-1} . The current is baroclinically unstable (e.g. Griffiths and Linden, 1981) and creates meanders which penetrate into the interior of the tank. During the forcing period of the experiment when the source continuously supplies fresh water, an entire surface of the tank eventually becomes covered with a layer of fresh water. Thus, two-layer system is created. The forcing then stops and the flow is allowed to develop freely. The system contains a large amount of available potential energy which is released gradually and maintains the flow for a very long time after the forcing stops. Measurements of barotropic and baroclinic components of velocity (for details see Matulka and Afanasyev, 2015) show that the upper layer rotates cyclonically while the lower layer rotates anticyclonically. The shear between the layers makes the system baroclinically unstable. Baroclinic instability continuously generates meanders over the entire area of the tank. The meanders move water parcels in the radial (meridional) direction. According to conservation of potential vorticity the parcels acquire additional relative vorticity and radiate Rossby waves. Motion of the meanders/parcels correlated via the global Rossby wave field creates the Reynolds stresses which drive zonal jets in the interior of the tank. Measurements of the Reynolds stresses showed that jets in the interior are dynamically different from the coastal jet which is affected by the presence of the wall. In what follows we perform spectral analyses of the flow in the inner area which contains these “true” jets and excludes the coastal current.

Visual comparison between the fields in the experiments with different forcing (Fig. 3) shows that the scales of the turbulent eddies generated by the forcing are noticeably different. The eddies in the thermal experiment are smaller which indicates smaller R_d . Another difference is perhaps more significant in distinguishing between the forcing

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mechanisms. The flow with saline forcing is characterised by thin filaments rather than circular eddies. In fact, eddies appear only as a result of breaking of the filaments and do not have very long lifetime. The filaments are created by the baroclinic instability; they protrude in the radial (meridional) direction. We hypothesise that these filaments are the manifestation of the so-called “noodles” which are the primary mode of the instability of the baroclinic flow (Berloff et al., 2009a, b).

3.2 Energy spectra in wavenumber space

Herein we describe the results of the spectral analysis of the flows. For a circular domain such as our tank, it is perhaps more natural to use polar coordinates for the purpose of spectral decomposition. Afanasyev and Wells (2005) used Fourier–Bessel transform to obtain two-dimensional energy spectra of the polar β -plane turbulence in the space of azimuthal and radial wavenumbers and then to obtain a one-dimensional spectrum by sorting data in terms of a polar analogue of the Cartesian isotropic wavenumber. However, it is easier to perform digital Fourier decomposition in the Cartesian coordinates because Fast Fourier Transform routines can be used. Moreover, usage of Cartesian coordinates and a conventional β -plane (rather than quadratic polar β -plane) simplifies the further comparison with available theory. For these reasons here we introduce a local Cartesian coordinates (x, y) centred at the reference radius $r_0 = 25$ cm such that x axis is directed to the east, $x = r_0\theta$ (where θ is the polar angle) and y axis is directed to the north, $y = r_0 - r$. For the spectral analyses we chose a domain of half width 17 cm centred at the reference radius $r_0 = 25$ cm such that the polar part of the tank, where β -plane approximation is inappropriate, and the wall area, where the coastal jet dominates, were excluded.

The AIV technique gives velocity field on a regular rectangular grid covering the entire area of the tank. The velocity vector field was interpolated onto the local Cartesian coordinate system and then projected to eastward and northward directions to obtain zonal and meridional velocity components. Discrete Fourier transform of these

velocity components then gives velocity $\mathbf{u}(k_x, k_y)$ in the wavenumber space (k_x, k_y) . Two-dimensional energy spectrum is given by

$$E(k_x, k_y) = \frac{1}{2} |\mathbf{u}(k_x, k_y)|^2. \quad (5)$$

Figure 4 shows two-dimensional spectra in the experiments with thermal and saline forcing. The spectra are measured in the beginning of both experiments when the spectral signature of the forcing is still strong, and at a later stage, when the spectra reached certain “saturation” and the turbulent cascades are developed. The spectral properties of forcing or background turbulent flow can be inferred from the initial spectra. In the experiment with thermal forcing the eddies generated by thermal plumes are quite small; their scale can be estimated from the value of wavenumber $k \approx 2.5 \text{ cm}^{-1}$ (outer ring in

Fig. 4a) to be approximately 2.5 cm. Here $k = (k_x^2 + k_y^2)^{1/2}$ is the isotropic wavenumber. These eddies are initially concentrated along the heating wires. The separation of the heating wires (approximately 4.5 cm) determines a wavenumber $k \approx 1.4 \text{ cm}^{-1}$ where energy concentration is observed as well (inner ring in Fig. 4a). In the experiment with saline forcing the initial spectrum is determined by the baroclinic instability which is sustained in the two-layer system even when the actual forcing (the injection of fresh water) is stopped. The energy is distributed in a wide range of wavenumbers but is mainly contained within a circle which corresponds to the reciprocal of the radius of deformation, $R_d^{-1} = 1.2 \text{ cm}^{-1}$. This is in agreement with the prediction of the Phillips model (Phillips, 1951) of baroclinic instability. The model predicts that all perturbations of wavenumber $k < R_d^{-1}$ are unstable with the most unstable (before nonlinear saturation) wavenumber being $k = 0.64 R_d^{-1}$. Thus the initial spectra for both experiments confirm our initial observation that the forcing scale is smaller in the experiment with the thermal forcing. The spectra measured in much later times in both experiments (Fig. 4b and d) show that energy cascades towards smaller wavenumbers (larger scales). The distribution of energy in the wavenumber space also becomes more anisotropic; the energy flows towards the k_y axis (zonal modes, $k_x = 0$).

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The inverse energy cascade is a well-known phenomenon in two-dimensional turbulence; energy is transferred from small scales (large k) to large scales (small k). In the presence of β -effect, this scenario is modified. The pioneering work by Rhines (1975) demonstrated that there is a certain scale, now known as the Rhines scale, that separates the large-scale motions where β -effect dominates from a small-scale turbulence. A wavenumber corresponding to the Rhines scale is given by

$$k_R = \sqrt{\beta/U_{\text{rms}}}, \quad (6)$$

where U_{rms} is root-mean-square velocity. On large scales Rossby wave elasticity is important and the flow becomes strongly anisotropic as the (linear) dispersion relation for Rossby waves suggests. The anisotropy is manifested by the appearance of zonal jets. The wavenumber k_R is also widely used as an estimate for the meridional wavenumber of an arrangement of zonal jets. It was shown to work well in different circumstances including flows on gas giants or flows in the laboratory although different modifications of the Rhines scale were discussed as well. The values of k_R are marked by crosses on y axis in Fig. 4b and d.

To extend the Rhines' argument to two dimensions one can equate the frequency of turbulent eddies to that of Rossby waves to obtain (Vallis and Maltrud, 1993) a dividing line in the form

$$k_\theta = \sqrt{\frac{\beta \cos \theta}{U_{\text{rms}}}}, \quad (7)$$

where θ is the polar angle in the wavenumber space, $\theta = \arctan(k_y/k_x)$. The line given by Eq. (7) resembles figure eight or a dumbbell and is shown in Fig. 4b and d. Note that values of U_{rms} are not given by theory but have to be measured in the experiment. Here we used the rms values of the y component of velocity instead of the total velocity in order to avoid the mean flow in the azimuthal direction which occurred to different extent in both of the experiments. Since Rhines' theory assumes isotropic small-scale

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character of the flow in the first experiment. Steeper slope in the second experiment is due to the particular nature of this two-layer baroclinic flow. Note that steeper than -3 slope was also measured in the experiments with shallow water, f-plane turbulence by Afanasyev and Craig (2013) and in the numerical simulations by Yuan and Hamilton (1994). The estimates of the Rhines wavenumber (accidentally) give similar values for both experiments, $k_R \approx 0.45 \text{ cm}^{-1}$. This value indicates the place for zonal jets in the spectral cascade. The maximum energy achieved by the cascade in the second experiment roughly corresponds to the Rhines wavenumber that indicates that zonal motions should possess a significant portion of the total energy. In the first experiment, the cascade does not quite reach the Rhines wavenumber. Note that the Rhines wavenumber does not stop the cascade to even lower wavenumbers; the cascade is just redirected towards zonal motions.

4 Conclusions

In our experiments we observed the formation of zonal jets in the experiments where flows were forced using two different methods. Perhaps the main difference between the forcing was that the heater at the bottom created convectively unstable vertical temperature distribution which resulted in small scale convective plumes. Vertical mixing must be significant in this system and the fluid remained mainly unstratified although its mean temperature increased during the forcing period. The large scale flow in this experiment is then approximately barotropic, although the nature of forcing is baroclinic. In the second experiment, on the other hand, we created statically stable two-layer stratification. The flow was baroclinic to a significant degree. Since this system was unstable with respect to baroclinic instability, the instability as a source of small-scale turbulence. Thus in both cases some small-scale turbulence was created but in the former experiment the flow was mainly barotropic while in the latter it was mainly baroclinic.



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In spite of this significant difference between the flows in our two experiments, we observed a certain universality in their spectral dynamics. The energy cascaded from small scales to larger scales and towards zonal motions. The two-dimensional spectra demonstrated that this cascade is in a reasonable agreement with the Rhines theory.

One-dimensional spectra of energy reliably demonstrated the existence of the energy interval with the $-5/3$ slope. Note that although in our experiments one can infer the direction of the energy cascade from the form of the spectrum assuming that the forcing wavenumber is known, direct evidence of the cascade direction can only be provided by the analyses of the spectral energy flux. Such evidence was provided for shallow water rotating turbulence (without β -effect) in the laboratory investigation by Afanasyev and Craig (2013) The analysis of the energy flux for the large-scale oceanic turbulence by Scott and Wang (2005) revealed the existence of the inverse cascade in agreement with the two-dimensional turbulence theory. However, the interplay between barotropic and baroclinic modes and the extent each mode contributes to the cascade in the ocean still remains a subject of research. The analysis of the flux for the experiments reported here is yet to be done and will be reported elsewhere.

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turbulence**

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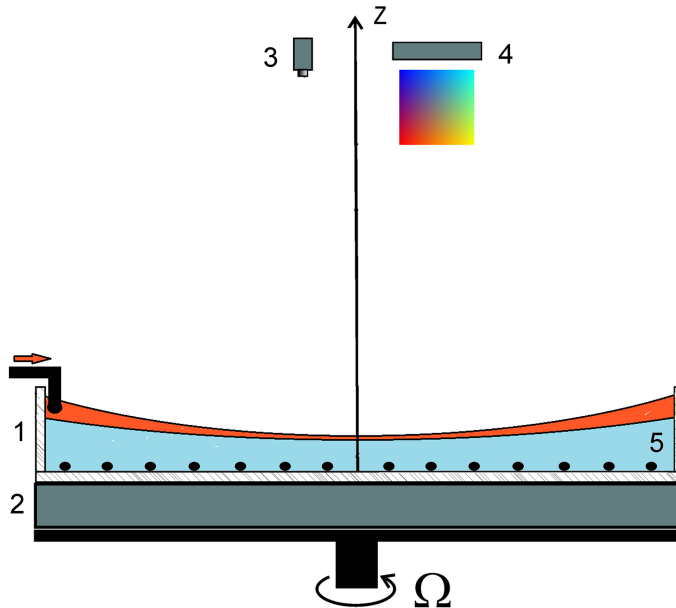


Figure 1. Sketch of the experimental setup: (1) cylindrical tank filled with water and installed on a table rotating with angular velocity Ω ; (2) light box for the optical thickness measurements; (3) video camera; (4) light source with color mask; and (5) heating wire on the bottom.

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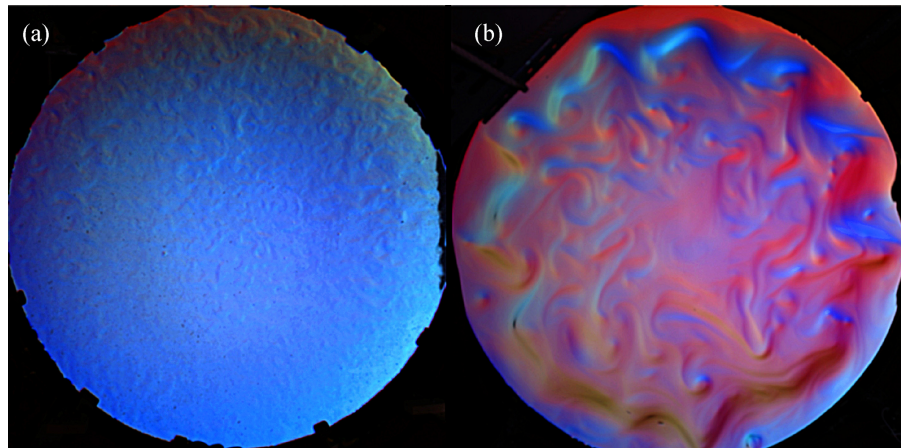


Figure 2. Typical images from video sequences recorded in the experiments with thermal **(a)** and saline **(b)** forcing. The flows are visualized by optical altimetry (AIV); different colors indicate different values (both in magnitude and direction) of the gradient of the surface elevation.

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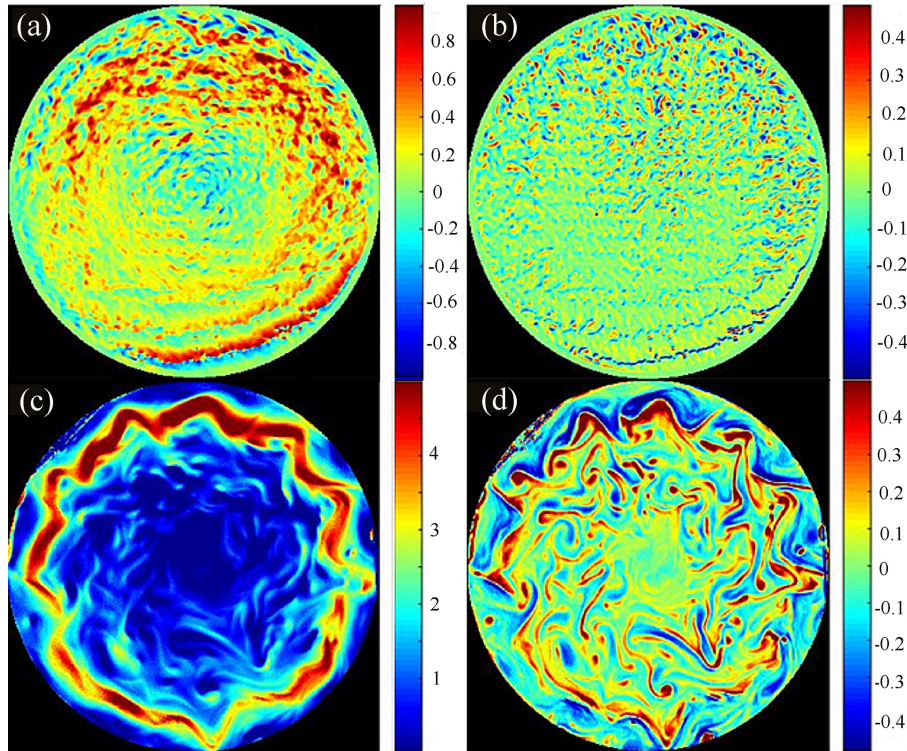


Figure 3. Flows generated by thermal forcing at $t = 280$ s (**a**, **b**) and by saline forcing at $t = 150$ s (**c**, **d**). (**a**) and (**c**) show the x component of velocity (azimuthal velocity), u , while (**b**) and (**d**) show the dimensionless relative vorticity, ζ/f_0 . Salinity difference between the layers for the experiment with saline forcing (**c**, **d**) is $S = 30$ ppt. The center of the tank corresponds to the North Pole of the polar β -plane.

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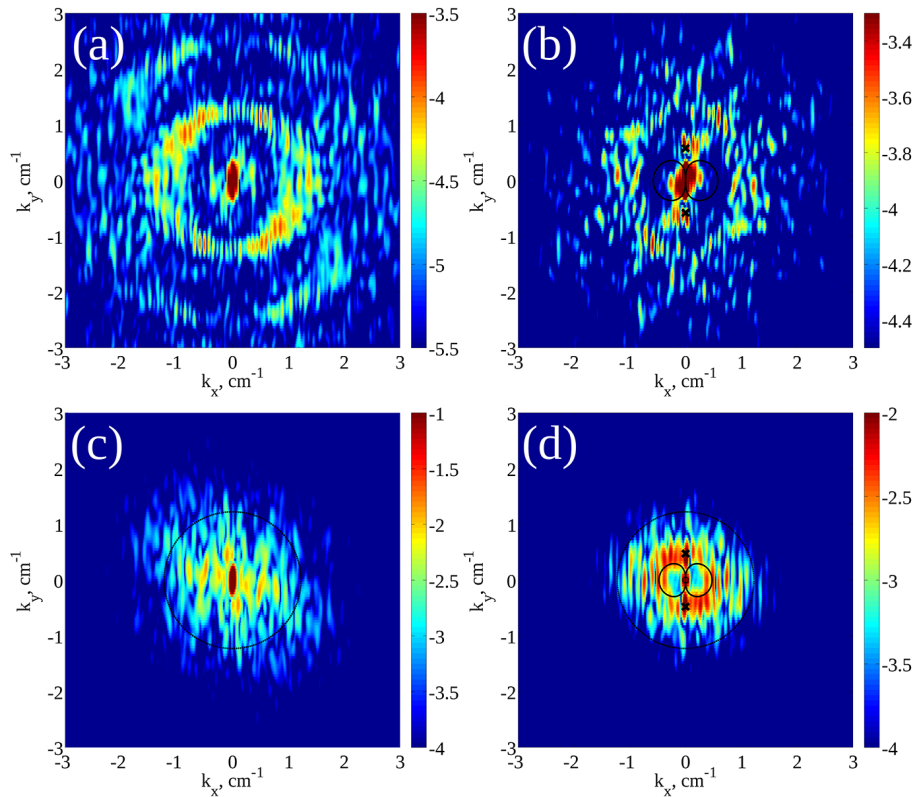


Figure 4. Energy spectra in the wavenumber space (k_x, k_y) for the experiments with thermal **(a, b)** and saline **(c, d)** forcing: $t = 30$ s **(a)**, 630 s **(b)** from the beginning of the thermal experiment; $t = 10$ s **(c)**, 722 s **(d)** after the fresh water source stopped in the experiment with saline forcing. Color shows energy in logarithmic scale. Black circles in **(c)** and **(d)** show R_d^{-1} . Black crosses in **(b)** and **(d)** represent the Rhines scale wavenumber (5) while the dumbbell-shaped curve is given by Eq. (7).

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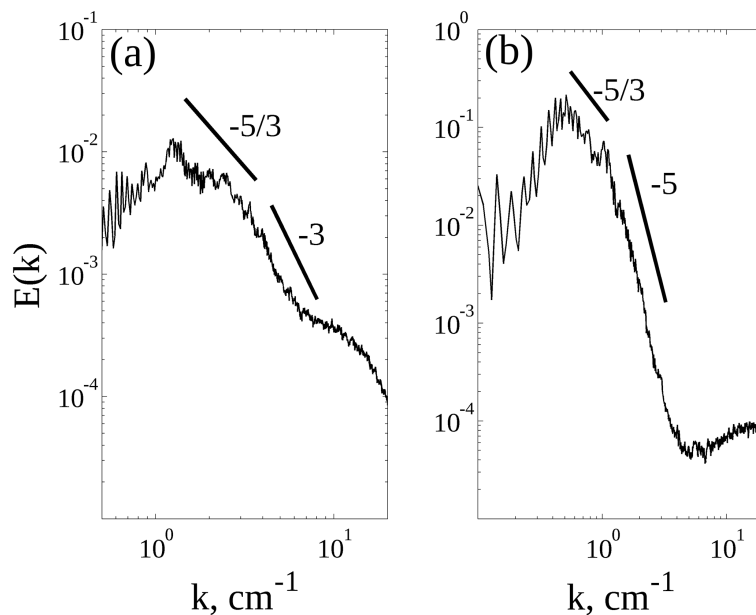


Figure 5. The one-dimensional energy spectra in log-log scale for the experiments with thermal (a) and saline (b) forcing: $t = 630$ s (a), 722 s (b) (as in Fig. 4b, d).

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