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Corrigendum to

"Time-varying magnetotail magnetic flux calculation: a test of the method" published in Ann. Geophys., 27, 1583–1591, 2009

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1. The solar wind dynamic pressure *Pd* was calculated as

   Pd=1.94×10^{-6}$n_{SW}V_{SW}^2$ (assuming 4%-helium content, e.g. Tsyganenko, 2002). (p. 1584)

   However, in the CCMC simulations there is no helium, and so Pd=1.67×10^{-6}$n_{SW}V_{SW}^2$, which slightly changes the result.

2. There is a misprint in Eq. (4) (p. 1585): instead of

   $$\sin^2\alpha = A\exp(CX) \quad (4)$$

   it should be

   $$\sin^2\alpha = A^2\exp(CX) \quad (4)$$

3. Taking into account all corrections, the final result is:

	X=−15 R_E	X=−25 R_E
Tail	F_{SPR}=0.91 F_D+0.26, cc=0.99	F_{SPR}=0.99 F_D+0.19, cc=0.94
lobes	F_{PR96}=0.70 F_D+0.35, cc=0.96	F_{PR96}=0.66 F_D+0.29, cc=0.94
beta=0.1	F_{SPR}=0.88 F_D+0.19, cc=0.79	
beta=0.5	F_{SPR}=0.83 F_D+0.21, cc=0.70	
beta=1	F_{SPR}=0.86 F_D+0.18, cc=0.70	
neutral sheet	F_{SPR}=0.7 1 F_D+0.20, cc=0.67	
	F_{PR96}=0.84 F_D+0.24, cc=0.75	

   The main conclusion of the work conserves: correspondence of F_{SPR} with F_D is good in the lobes and spoils when approaching the neutral sheet, though the regression equation is rather stable. In the tail lobes the SPR algorithm gives better results, than PR96, and vice versa in the neutral sheet, which is probably due to one-dimensional pressure balance violation.

M. A. S. would like to apologize for her inaccuracy.

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