

Transfer of organic Br and Cl from the Biosphere to the Atmosphere during the Cretaceous/Tertiary Impact: Implications for the stratospheric Ozone Layer

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Abstract. Following the Cretaceous/Tertiary (K/T) meteoritic impact some 65 Myr ago, large portions of aboveground terrestrial biomass were burned. As a result, large amounts of various trace gases were injected to the atmosphere, inducing a wide range of effects on climate and ecosystems. Here, it is commented on the previously unaccounted for emission to the atmosphere of methyl bromide (CH_3Br) and methyl chloride (CH_3Cl) from extensive biomass burning that followed the impact. Based on reported biomass burning emission rates of the above organohalogen relative to CO_2 , it is estimated that their emissions from global fires resulted in tropospheric mixing ratios of around 20–65.8 ppbv organic Cl and 110–390 pptv organic Br. The above calculated mixing ratios of organic chlorine and bromine are more than an order of magnitude greater than their present, anthropogenically perturbed level and, although the ocean ultimately might absorb them, we argue here that they could still remain in the atmosphere for many years, and a substantial fraction could be transported to the stratosphere, thus substantially affecting the ozone layer. This would have led to very serious increases in short wavelength UV radiation reaching the lowermost atmosphere.

1 Introduction

Atmospheric consequences of meteoritic impacts have been discussed by a number of researchers (see review by Toon et al., 1997). These consequences include climatic and chemical effects in the troposphere and the stratosphere. Massive impacts, with energies $> 10^7$ Mt, occur at intervals of approximately 10^7 yrs and the ejecta re-entering the atmosphere can lead to massive global fires that can burn a major part of the terrestrial aboveground biomass. Reported stratospheric con-

sequences of bolides $> 10^7$ Mt result from injections of dust, NO, and water vapour in the stratosphere. Ozone loss is then a consequence of enhanced NO levels and accelerated heterogeneous chemistry. After the removal of radiation-blocking dust, smoke, and NO_2 from the atmosphere, significantly increased doses of UV-B radiation could reach the ground for periods up to several years as a result of ozone depletion. However, if a shielding layer of stratospheric sulphate particles forms, much of the UV-B may be blocked from reaching the surface during the first few years after the impact.

Tropospheric effects might be even more serious. During the K/T boundary event some 65 Myr ago, which is currently believed to have been caused by an object with impact energy of the order of 10^8 – 10^9 Mt, 3×10^{17} – 3×10^{18} g NO might have been produced. These NO levels could cause mass extinctions through a series of effects, the most serious of these being extensive acidification of surface soils and waters due to the precipitation of nitric acid (HNO_3), produced by the atmospheric degradation of NO (Prinn and Fegley, 1987; Crutzen, 1987).

It is believed that a major fraction of the aboveground biomass burned after the K/T impact from fires ignited by ejecta re-entering the atmosphere (Kring and Durda, 2002). This is because of a worldwide soot layer at the K/T boundary (Wolbach et al., 1990a, b), carbon isotope shifts (Ivany and Salawitch, 1993), and a widespread, sharp peak in the distribution of fern spores (Wolfe and Upchurch, 1986), the latter usually taken as a sign of wildfires. Extrapolating from the discovery of 5×10^6 g of soot in K/T clays (Wolbach et al., 1990a, b), Crutzen (1987) estimated that global fires following the impact produced more than 10^{19} g CO_2 , 10^{18} g CO, and a variety of other gases, including pyrotoxins.

Methyl bromide and methyl chloride have not yet been included in the calculated trace gas emissions to the atmosphere during the K/T impact. However, these gases are sufficiently long-lived to enter the stratosphere, where they

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release active chlorine and bromine, and thus catalytically destroy ozone. Here, recent reports of CH_3Br and CH_3Cl emissions from biomass burning are used to calculate the amounts of these gases that would have been released from the global fires of the K/T impact event.

2 Results and discussion

The halogenated gases CH_3Cl and CH_3Br deliver Cl and Br, respectively, to the stratosphere, where they can catalytically destroy ozone (e.g. Fabian et al., 1994; Kourtidis et al., 1998; Montzka and Fraser, 2003). CH_3Cl is the most abundant halocarbon in the atmosphere, with a NH (SH) average mixing ratio around 536 (541) pptv (Simmonds et al., 2004), corresponding to an atmospheric global burden of about 4.3 Tg. Biomass burning might presently be the single largest source of methyl chloride ($1.0 \pm 0.3 \times 10^{12} \text{ g yr}^{-1}$, Andreae et al., 1996; Blake et al., 1996; Lobert et al., 1999), followed by oceanic emissions ($0.3 \pm 0.1 \times 10^{12} \text{ g yr}^{-1}$, Moore et al., 1996; $0.65 \times 10^{12} \text{ g yr}^{-1}$, Khalil et al., 1999) and fungal production ($0.16 \times 10^{12} \text{ g yr}^{-1}$, Watling and Harper, 1998). The amounts of methyl halides produced during biomass burning are related to the Cl/Br content of the biomass, which again depends on the soil Cl/Br content (Andreae et al., 1996; Gan et al., 1998). Sinks of methyl chloride are OH attack in the troposphere (-3 Tg yr^{-1}) and, to a lesser extent, oceanic uptake (-0.2 Tg yr^{-1}). CH_3Br exists currently in the atmosphere at a NH (SH) mixing ratio of 10.4 (7.9) pptv (Simmonds et al., 2004), corresponding to an atmospheric global burden of about 146 Gg. Manö and Andreae (1994) first reported biomass burning as a source of methyl bromide. This source delivers currently $10\text{--}24 \text{ Gg yr}^{-1}$ (best estimate 20 Gg yr^{-1}) CH_3Br to the atmosphere and is thus responsible for around 20% of total global emissions of CH_3Br (Andreae et al., 1996; Blake et al., 1996). Sinks of methyl bromide other than the oceans include reaction with OH and photolysis (-86 Gg yr^{-1}) and uptake by soils (-42 Gg yr^{-1} , Shorter et al., 1995). Not all source terms of CH_3Cl and CH_3Br are included above, hence the large apparent imbalance between source and sink terms. However, the reader should keep in mind, that even with all best estimate source/sink terms included, the budget of both CH_3Cl and CH_3Br would still remain imbalanced, by -1.05 Tg yr^{-1} and -45 Gg yr^{-1} , respectively. If the full range of the source/sink estimates is considered, the budgets end up with a net imbalance range -4.6 to $+9.3 \text{ Tg yr}^{-1}$ for CH_3Cl and -220 to $+71 \text{ Gg yr}^{-1}$ for CH_3Br (Montzka and Fraser, 2003). The imbalance most probably results from uncertainties in current source/sink estimates and/or unidentified source terms. Montzka and Fraser (2003) provide an up-to-date review of the subject.

Andreae et al. (1996) during the SAFARI-92 experiment and Blake et al. (1996) during the TRACE-A experiment have measured the relative emission ratios of methyl halides. Andreae et al. (1996) report emission ratios of CH_3Cl rela-

tive to CO_2 and CO of 20 ppt/ppm and 0.95 ppt/ppb, respectively, and emission ratios of CH_3Br relative to CO_2 , CO and CH_3Cl of 0.11 ppt/ppm, 8.3×10^{-3} ppt/ppb, and 8.3 ppt/ppb, respectively. Blake et al. (1996) measured emission ratios of CH_3Cl relative to CO_2 and CO of 27–37 ppt/ppm and 0.57–0.85 ppt/ppb, respectively, and emission ratios of CH_3Br relative to CO_2 and CO of 0.37–0.39 ppt/ppm and $6\text{--}10.6 \times 10^{-3}$ ppt/ppb. Lobert et al. (1999), after a compilation of all available data, arrived at a significantly higher median emission ratio of CH_3Cl relative to CO_2 of 65.8 ppt/ppm. Hence pyrogenic CH_3Br and CH_3Cl emissions are today considered as significant contributors to stratospheric ozone depletion, constituting roughly 25% and 20% of present-day global CH_3Cl and CH_3Br emissions, respectively (Andreae et al., 1996; Blake et al., 1996). Other authors also reported relative emission ratios of CH_3Cl to CO and CO_2 (see compilation by Lobert et al., 1999, and references therein). As noted above, shortly after the K/T impact, around 10^{19} g of CO_2 have been emitted to the atmosphere from extensive global fires, about 3 times the amount of CO_2 currently in the atmosphere, adding globally around 1000 ppm CO_2 to the troposphere (Crutzen, 1987).

Hence, by using the abovementioned emission rates relative to CO_2 , we calculate that the amounts of CH_3Cl and CH_3Br that were released into the atmosphere would have resulted in global tropospheric volume mixing ratios (VMRs) of around 20–65.8 ppbv of CH_3Cl and 110–390 pptv CH_3Br after the K/T impact. The above-calculated VMRs of methyl chloride are 36–120 times its present level of about 550 pptv, while the methyl bromide VMR is 11–39 times its present level of about 10 pptv. Considering that once activated, the impact-induced halogen loadings translate to 20–65.8 ppbv of active chlorine and 110–390 pptv of active bromine, and that the present-day, anthropogenically perturbed levels of organic chlorine and bromine are around 3.6 ppbv and 20 pptv, respectively (Montzka and Fraser, 2003), the impact-induced loadings would have seriously damaged the stratospheric ozone layer, since they are 5.6–18.3 times and 5.5–19.5 times, respectively, of their present-day levels.

Whether such damage to the ozone layer would have occurred or not, ultimately would depend on:

1. The efficiency of tropospheric removal through reactions with OH, if large amounts of CH_3Cl and CH_3Br do not reach the stratosphere soon after the impact. The tropospheric lifetime of these two gases is presently determined by the hydroxyl radical (OH) abundance. Due to the burning of substantial portions of aboveground biomass, concurrently with the release of CH_3Cl and CH_3Br , substantial amounts of smoke and radiation-absorbing gases (like NO_2) would have been emitted to the atmosphere. Smoke and absorbing gases will block solar radiation, and OH production from the photodissociation of tropospheric ozone will either stop or be reduced substantially. This, combined with fast

consumption of OH by emitted hydrocarbons, will lead to OH levels near zero and hence very long tropospheric lifetimes for CH₃Cl and CH₃Br until the smoke veil disappears.

- The efficiencies of tropospheric removal through reactions on particles, if large amounts of CH₃Cl and CH₃Br do not reach the stratosphere soon after the impact. Heterogeneous tropospheric and stratospheric chemistry after such an explosion would be greatly perturbed (see review by Toon et al., 1997). It cannot be said with certainty whether the removal of the above gases from the atmosphere through heterogeneous reactions on tropospheric particles would substantially deplete them since, to our knowledge, no kinetic data exist that would allow for a quantification of this effect. However, since most published data on relative emission rates of these methyl halides from biomass burning are from in-situ observations of plumes, young or aged, uptake on particles, if any, is likely already included in our calculations.
- The efficiencies of tropospheric removal through oceanic uptake, if large amounts of CH₃Cl and CH₃Br do not reach the stratosphere soon after the impact. How fast and whether oceanic uptake would reduce significantly the atmospheric burden would depend on a number of factors, related to the biological, chemical and thermal state of the ocean after the impact, some of which are poorly constrained. Namely, a) we do not know whether the massive die-off of phytoplankton that occurred after the impact (e.g. Kaiho et al., 1999) would have released copious amounts of intracellular methyl bromide and methyl chloride in the sea water; this would temporarily reduce the degree of ocean undersaturation, or might even have resulted in supersaturation with respect to the atmosphere, and b) we do not know the degree of inhibition of vertical mixing after the impact; if vertical mixing was seriously reduced, stratification of the upper ocean layer would inhibit the loss of surface water organohalogenes to the ocean below, resulting in quick saturation of the upper layer. Here, we use a two-box ocean-atmosphere model and some reasonable assumptions to calculate the effect of ocean uptake. Due to the absence of a relevant model for CH₃Cl, only CH₃Br was modelled.

The model is identical to that in Butler (1994). Ignoring transport, the general equations governing the coupled ocean-atmosphere system are (Butler, 1994; Yvon and Butler, 1996):

$$\frac{dn_a}{dt} = F_{oa} - F_{OH}, \quad (1)$$

$$\frac{dn_o}{dt} = P_o - F_{oa} - F_d - F_{ed}, \quad (2)$$

where n_a and n_o are the mole mass of CH₃Br in the atmosphere (equal to $p_a M$, where p_a is the atmospheric partial pressure of CH₃Br at sea level in patm and $M=1.79 \cdot 10^{20}$ mol is the mass of the atmosphere) and in the surface ocean (equal to $p_w A z / H$), respectively, F_{oa} is the net flux from the ocean to the atmosphere (mol·yr⁻¹), F_a is the atmospheric loss rate of CH₃Br (mol·yr⁻¹), which equals the loss rate due to reaction with OH, F_{OH} , P_o is the production of CH₃Br in the ocean (mol·yr⁻¹) ($P_o = p_o \cdot A$, where p_o is the production rate in the ocean in mol·m⁻²·yr⁻¹ and A is the entire surface area of the ocean and equal to $3.61 \cdot 10^{14}$ m²), F_d is the aquatic degradation rate (mol·yr⁻¹), and F_{ed} is the loss rate of CH₃Br from combined downward mixing and degradation through the thermocline (mol·yr⁻¹).

The analytical expressions of Eqs. (1) and (2) are

$$\frac{dn_a}{dt} = \frac{K_w A}{H} (p_w - p_a) - k_{OH} n_a \quad (3)$$

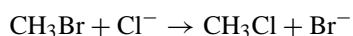
$$\frac{dn_o}{dt} = P_o - \frac{K_w A}{H} (p_w - p_a) - k_d n_o - \frac{\sqrt{D_z k_d}}{z} n_o, \quad (4)$$

where k_d is the pseudo-first order loss constant for aquatic degradation in the surface layer (36 yr⁻¹), K_w is the mean air-sea exchange coefficient ($1.5 \cdot 10^3$ m·yr⁻¹), p_a is the partial pressure of CH₃Br (atm) in the atmosphere at sea level and p_w is the mean oceanic partial pressure of CH₃Br in patm, H is the Henry's Law constant (m³·atm·mol⁻¹), k_{ed} is the "eddy degradation" rate for removal of CH₃Br downward from the surface layer (440 m·yr⁻¹), z is the thickness of the surface layer (m), assumed to be constant and equal to 75 m, k_a is the pseudo-first order loss constant for degradation in the atmosphere, and if we treat only the reaction with OH k_a equals k_{OH} (the CH₃Br stratospheric loss has a rate which is slow compared to OH and is ignored here). The eddy diffusion rate of removal, k_{ed} , is given by

$$k_{ed} = (D_z k_d)^{0.5}, \quad (5)$$

where the mean diffusivity through the thermocline, D_z , is equal to 5440 m·yr⁻¹.

The most significant tropospheric removal pathway for CH₃Br is the reaction with the hydroxyl (OH) radical, which determines the lifetime of CH₃Br in the atmosphere. A mean OH tropospheric concentration of 10^6 molecules cm⁻³ and a reaction rate of $k_{OH} = 1.86 \cdot 10^{-12} \cdot e^{-1230/T} \text{cm}^3 \cdot \text{molecule}^{-1} \cdot \text{s}^{-1}$ were used (Chichinin et al., 1994). The main aquatic chemical degradation pathways for CH₃Br are nucleophilic substitution by Cl⁻ and hydrolysis, the former being the main destruction pathway,



at a typical oceanic surface temperature of 21.9°C, [Cl⁻]=0.56 mol·liter⁻¹. For $k_{\text{H}_2\text{O}}$ and k_{Cl} , the values by Jeffers and

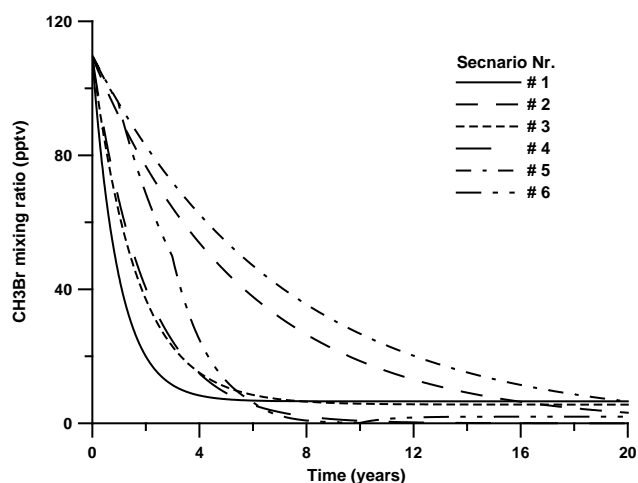


Fig. 1. Box-model calculations, for different scenarios (see Table 1), of the evolution of the tropospheric burden of CH_3Br after its initial release from biomass burning. The initial tropospheric molar ratio of CH_3Br was set at 110 pptv, which is a lower limit for the initial release.

Wolfe (1996) were used, $k_{\text{H}_2\text{O}} = 5.7 \cdot 10^{11} \cdot e^{-(12455 \pm 240)/T} \text{ s}^{-1}$, $k_{\text{Cl}^-} = 1.09 \cdot 10^{13} \cdot e^{-(12724 \pm 560)/T} \text{ l mol}^{-1} \cdot \text{s}^{-1}$.

Equations (3) and (4) were solved numerically by Euler's method. As the model is formulated, CH_3Br is produced solely in the ocean. In the baseline run, we used a P_0 value of $6.0 \cdot 10^{-6} \text{ mol m}^{-2} \text{ yr}^{-1}$, half of the sum of P_0 in the open ocean, $3.9 \cdot 10^{-6} \text{ mol m}^{-2} \text{ yr}^{-1}$, and the coastal water, $8.0 \cdot 10^{-6} \text{ mol m}^{-2} \text{ yr}^{-1}$, as given by Butler and Rodriguez (1996). We also assume that the removal of atmospheric CH_3Br takes place only by its reaction with OH and by ocean-atmosphere exchange. Depending on the initial conditions, the latter process is a sink or a source of atmospheric CH_3Br . The soil/canopy sink as well as newly discovered sources from fungi, coastal salt marshes, plants, wetlands and natural oxidation processes during degradation of Fe-containing organic matter were not considered, since they are today less important than the ocean in defining the budget of methyl bromide and were probably even less important in the perturbed conditions after the K/T impact. Possible oceanic and atmospheric heterogeneities, such as the spatiotemporal variability of mixing depth z and Cl^- and OH concentrations, were ignored.

To investigate the possible development of tropospheric CH_3Br in time after the initial emission from the massive impact fires, different scenarios were modelled (Table 1 and Fig. 1). These scenarios can give insight into the possible development of CH_3Br concentrations during post-impact conditions. Although only the lower limit of initial releases was modelled (110 pptv), for other concentrations the relative development will be rather similar.

Under present-day, baseline conditions (Scenario 1), CH_3Br is relatively quickly removed from the atmosphere

due to OH oxidation. Under perturbed conditions (Scenarios 2 to 6), although we assumed that oceanic production ceases or is greatly diminished, CH_3Br is removed more slowly, due to the assumed reductions in the oxidation capacity of the atmosphere. If atmospheric oxidation ceases completely for the simulation period (Scenario 2), around 1/3 of CH_3Br still remains in the atmosphere after 6 years. This is an improbable scenario, corresponding to the persistence of a heavy smoke and dust cloud, that demonstrates the strong influence of the oxidising capacity of the atmosphere on the post-impact development of tropospheric CH_3Br .

Scenarios 3 and 4 show quite similar developments during the first 4 years. For these scenarios, where oceanic production of CH_3Br ceases or is halved and atmospheric oxidation is halved concurrently with changes in mixed layer depth, around 1/3 of CH_3Br still remains in the atmosphere after 2 years.

If we assume that the atmospheric oxidation and the oceanic production cease, the mixed layer depth is halved, and at the same time large amounts of intracellular CH_3Br are released in the seawater due to the massive death of plankton (Scenario 5), then tropospheric CH_3Br decreases even more slowly, due to the decrease of its flux across the air-sea interface.

Finally, Scenario 6 models a more probable development. We assume that oceanic production ceases for 10 years, and then slowly recovers at 30% its present-day value. The oxidising capacity of the atmosphere (as determined from OH concentrations) ceases for the first year due to the large amounts of smoke that block incoming radiation, and then slowly recovers as the smoke settles (see Table 1). We also assume that the ocean is more stratified and the mixed layer depth is halved, although this latter assumption will not have a major effect in the development of tropospheric CH_3Br , as simulation 4 has shown. For this scenario, more than half of the initially released CH_3Br still remains in the atmosphere after 2 years; after 4 years, about 1/4 is still in the atmosphere.

Hence, although ultimately the initially released tropospheric CH_3Br will be removed due to oxidation and absorption by the ocean, this removal will take some years to occur. It is important, however, to stress that damage to the ozone layer would ultimately depend largely on the efficiency of a fourth process, namely direct transport of the emitted amounts to the stratosphere. Transport to the stratosphere would certainly be influenced by the perturbed thermal state of the atmosphere. It has been argued, that fire-convection can entrain trace gases and smoke to the stratosphere (Fromm et al., 2004). Kao et al. (1990) modelled post-nuclear war conditions taking into account solar heating of the smoke plume. In their treatment, because very little sunlight reaches the surface the first days after the nuclear blasts, a thermal inversion occurs which prevents downward diffusion of heat and, additionally, sunlight heats the top of the smoke cloud making the atmosphere above unstable, thus

Table 1. The different model scenarios. P_o is the CH_3Br biological production in the surface ocean, and z is the mixed layer depth. The temperature of the oceanic surface layer, T_o , was set to 288°K, and the atmospheric temperature, T_a , was set to 290°K. The initial tropospheric molar ratio of CH_3Br was set at 110 pptv.

Scen. Nr.	P_o ($6 \cdot 10^{-6} \text{ mol m}^{-2} \text{ yr}^{-1}$)	Sea conc. ($\text{pmol} \cdot \text{mol}^{-1}$)	$[\text{OH}] (10^6 \text{ molecules} \cdot \text{cm}^{-3} \cdot 10^6)$	z (75 m)	Case
1	1	5	1	1	Baseline (present-day conditions)
2	0	5	0	1	Oceanic production and atmospheric oxidation cease
3	0.5	5	0.5	1	Oceanic production and atmospheric oxidation are halved
4	0	5	0.5	–	Oceanic production ceases, atmospheric oxidation is halved and there is no loss through the thermocline
5	0	50	0	0.5	Oceanic production and atmospheric oxidation cease, the mixing layer depth is halved and the sea conc. is raised due to release of intracellular CH_3Br
6	0 for 10 yr, 0.3 afterwards	5	0 until yr 1, 0.2 until yr 3, 0.6 until yr 6, 1 afterwards	0.5 until yr 1, 0.7 until yr 3, 1 afterwards	Mixed scenario

lofting the plume (Kao et al., 1990). Indeed, smoke from California fires in 1987 has been observed to strengthen a valley inversion cooling the surface by several degrees (Robock, 1988). Recently, smoke from fire plumes has been detected in the lower stratosphere (Fromm et al., 2000) and carbon monoxide and particles unambiguously attributed to North American fires were detected several kilometres above the extratropical tropopause (Jost et al., 2004). Proposed mechanisms responsible for the transport include convective systems inertial overshooting of the neutral buoyancy level, additional radiative self-heating of the plumes and subsequent lofting (Jost et al., 2004), as well as solar-driven lofting of smoke-filled air (Kao et al., 1990). Kring and Durda (2002) have calculated that the power delivered to the atmosphere by re-entering high-energy ejecta would have been in some areas in excess of 100 kWm^{-2} , with a peak of around 350 kWm^{-2} (see Kring and Durda, 2002, for details of the simulations), while the power delivered to the ground would be in excess of 12.5 kWm^{-2} for >20 min (the amount of energy sufficient to ignite vegetation) in large parts of the globe. The dissipation of such large amounts of energy, both from the re-entering ejecta and the subsequent global fires would probably involve also intensification of vertical mix-

ing and might also be sufficient to cause tropospheric folding in some areas and maybe even complete break-up of the tropopause barrier. Such Stratosphere-Troposphere Exchange (STE) events would bring large amounts of the emitted gases from the troposphere to the stratosphere, although it is not possible to quantify the exact extent of such STE events.

It thus appears very plausible that a large amount of the emitted methyl halides could reach the stratosphere, causing serious damage to the ozone layer for the following decades. The resulting ozone depletion would have influenced the transfer of solar UV through the atmosphere. Increases in UV would not reach the ground before tropospheric soot particles settle, as much of the radiation would be scattered by the particles with an optical depth of around 10^3 – 10^4 (see review by Toon et al., 1997). Spherical particles with a density of 3 g cm^{-3} falling in a Stokes regime from 85 km altitude would require more than a decade to settle if they have a radius of $1 \mu\text{m}$ and an hour if they have a radius of 1 mm (H. J. Melosh, cited in Fig. 12 of Kring and Durda, 2002). These are upper time limits, since coagulation of particles into larger grains was not considered. Other calculations suggest that a significant portion of sub-micron dust would have

remained in the atmosphere for several months, causing reductions in the light levels reaching the ground (e.g. Covey et al., 1990). It is unlikely that impact dust would cause much reduction in the UV part of the spectrum, except perhaps for the initial phase, when it would be denser. This is because dust causes attenuation mainly in the visible and UV-A part of the spectrum, while the shorter UV-B wavelengths remain unaffected (e.g. Madronich, 1993), as has been shown, for example, during Saharan dust events (Balis et al., 2002).

Ozone loss during such impacts is not presently considered the major threat to the biosphere, mainly because it has been assumed that ozone depletion would result from NO injections to the stratosphere. These would perturb the ozone layer, but any increase in UV radiation reaching the ground would be prevented by the shielding from smoke, dust, absorbing gases and a stratospheric sulphate particle layer that would form from the injection of large amounts of SO₂. Smoke and dust would settle after a few months, and the sulphate layer would also settle after a few years. Following the large eruption of volcanoes such as Pinatubo, the sulphate layer that formed had a persistence of about 2 years. After that period, significant doses of UV might reach the ground, but this effect would not last long since NO would also be removed from the stratosphere through conversion to HNO₃ and absorption on particles.

However, the calculations and argumentation presented in this manuscript imply that the ozone layer could be depleted considerably by the release of large amounts of relatively long-lived CH₃Br and CH₃Cl in a stratosphere already perturbed by large amounts of NO and particles. Such injections could raise the atmospheric abundance of natural organic chlorine and bromine an order of magnitude above their levels in the present-day, CFC and Halon-containing atmosphere. Ozone depletion would take place during the first years after such an impact in a stratosphere heavily loaded with aerosols and NO, the former speeding up the activation of Cl and Br. CH₃Cl and CH₃Br have presently an overall atmospheric lifetime of around 1.5–2 years. Because the inverse overall lifetime is the sum of the inverse stratospheric and tropospheric lifetimes, their overall lifetimes are primarily determined by their short (1.5–2 years) tropospheric lifetime, while their stratospheric lifetimes are considerably longer. CH₃Cl has a stratospheric lifetime of around 50 years, while CH₃Br has a somewhat lower stratospheric lifetime. The stratospheric lifetimes of the two species are quite long, compared with the lifetime of stratospheric aerosols of around one year and a lifetime for tropospheric aerosols of the order of a few days to a few weeks, depending on their size. The consequences of these organohalogen releases would be more persistent than aerosol injections and hence radiation enhancements due to the ozone depletion from natural organohalogen releases would outlive the tropospheric aerosol layer and the stratospheric sulphuric aerosol layer. Once the UV-B shielding of the aerosol layer ceased, considerable enhancements in short-wavelength UV radiation reaching the

ground could then persist for decades. Hence, ozone loss appears to be a major threat to the post-K/T impact biosphere, although the major threat could still be the fires and the resulting pyrotoxins and highly acidic wet and dry deposition preceding the high UV levels due to halogen-induced ozone loss. Recent work favours a role of UV in biological evolution (e.g. Cockell, 2000), although calculations by Cockell and Blaustein (2000) for K/T impact UV-B increases underestimate the role of increased UVB for exacerbating the demise of land animals due to the assumption of only NO-related ozone depletion. On the basis of the present calculations, it can be argued that the high post-impact UV levels implied by the present work would affect the species that survived the other two threats, and, perhaps most importantly, the post-impact recovery of the biosphere and the resulting species composition during the long-term recovery.

3 Conclusions

The results from the present calculations indicate that following the K/T impact, methyl bromide and methyl chloride emissions from biomass burning were of a magnitude that could seriously damage the ozone layer and as a consequence increase the penetration of short-wavelength solar radiation. A better study of the processes controlling the oceanic fluxes of these compounds and a better knowledge of the thermal and chemical state of the atmosphere and the ocean following the impact could offer additional constraints to the calculations presented here.

While impacts such as the K/T event occur at intervals of 10⁷ yrs, more frequent, less massive events could ignite a smaller portion of global forests. If one-tenth of the forests ignited during the K/T event are burned, CH₃Cl and CH₃Br levels might rise by 3.6–12 times and 1.1–3.9 times their present level, respectively. Global nuclear war could also cause extensive forest burning, although one should then bear in mind that present-day above-ground biomass density is around 1/3 of that of the end-Cretaceous period. It appears that pyrogenic emissions of these organohalogenes might also in such cases influence the ozone layer and should be included in relevant simulations. Lesser impacts would not perturb thermally the atmosphere to the extent of the K/T event; they would also not cause massive plankton die-off; would not alter in such dramatic ways the oxidizing capacity of the atmosphere and would not cause such changes in the global structure of the oceans. This might make such impacts more straightforward to model, since the range of the assumptions and uncertainties is much smaller and hence the involved parameters in the air-sea transfer calculations can be better constrained.

Only very recently some attention was given to the possible role of organohalogenes in earth history (Yokouchi et al., 2002). The present study shows that there is more reason for the study of these gases in such a perspective.

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